

"The original is unfaithful to the translation":
 (Kinetic theory, from Planck to galactice) ^R → C. Hamilton + J.B. Founry,
 P&P review

What is a galaxy? A collection of many stars w/ $\frac{1}{r}$ pot.

What is a plasma? A collection of many charges w/ $\frac{1}{r^2}$ pot.

To describe these,

$$\frac{d\vec{R}_i}{dt} = \frac{\partial H}{\partial \vec{V}_i} \quad \& \quad \frac{d\vec{V}_i}{dt} = -\frac{\partial H}{\partial \vec{R}_i} \quad \text{for each } i=1\dots N$$

\downarrow \downarrow
 \vec{V}_i $-\vec{\nabla} \Phi_g \text{ or } q \alpha \left(\vec{E} + \vec{V}_i \times \vec{B} \right)$

Compute electromagnetic fields produced by the charged particles, or gravitational fields produced by the masses, then use these fields to evolve the particles' phase-space positions. Done! Well...

- ① $\sim 10^{28}$ molecules in this room. Double-precision data dump would be $\sim 10^{18}$ TB! ($\sim 10^{71}$ protons in a cluster of galaxies)
- ② Not really interested in every single particle.
- ③ Many-body systems are often chaotic — initial conditions are arbitrary, and infinitesimal changes to them yield microscopically different results.

→ Statistical approach. What is prob. of a particle having \vec{r} and \vec{v} in some 6D interval $d^3\vec{r}d^3\vec{v}$?
 How does this probability evolve? Under what conditions is this probabilistic evolution accurate enough to yield reliable predictions for a single realization of the system?
 "We can measure the globula of matter and the distances bet. them, but Space itself is incomputable." — Nabokov

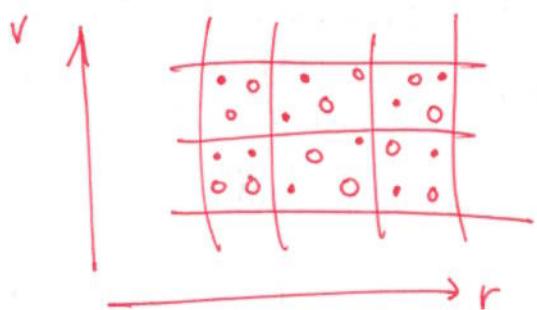
→ KINETIC THEORY.

This works because there is no difference statistically bet. flipping a coin $\sim 10^{28}$ times and flipping ~~one~~ $\sim 10^{28}$ coins one time each. Will say later why this is good for plannas... Adag or Ardor.

Start w/ Boltzmann distribution:

$$F_\alpha(t, \vec{r}, \vec{v}) = \sum_{i=1}^{N_\alpha} \delta(\vec{r} - \vec{r}_{\alpha i}(t)) \delta(\vec{v} - \vec{v}_{\alpha i}(t))$$

Species label



$$(N_\alpha = \int d\vec{r} \int d\vec{v} F_\alpha)$$

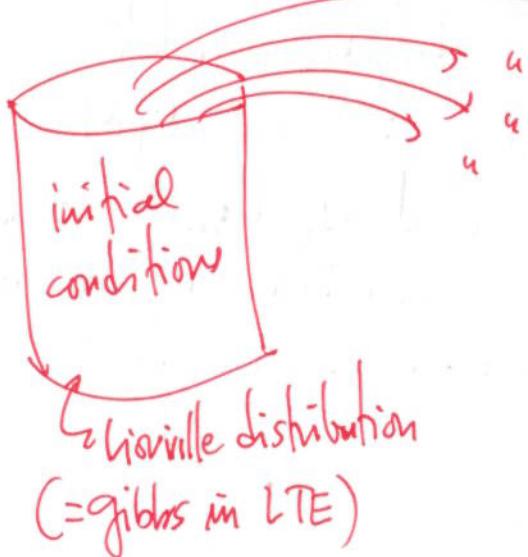
Can show that

$$\frac{\partial}{\partial t} \left(\frac{\partial F_\alpha}{\partial t} + \vec{v} \cdot \vec{\nabla}_r F_\alpha + \frac{\partial F_\alpha}{\partial v} \left(\vec{E}_m + \frac{\vec{v}}{c} \times \vec{B}_m \right) \right) = 0$$

(a Liouville thm. for point particles)

$$[F_\alpha, H_m]$$

Trade $6N$ eqns. to 1 eqn. with $6N$ dimensions.
It's still deterministic and nothing lost. So...
run through Klimontovich



Average the deterministic outcomes:

$$\left\langle \frac{\delta F_\alpha}{\delta t} + [F_\alpha, H_m] \right\rangle = 0 \quad P_N$$

What is $\langle F_\alpha \rangle_{P_N}$? "one-particle dist. function": prob. that particle α , has coordinates in \vec{X}_α to $\vec{X}_{\alpha_1} + d\vec{X}_\alpha$, regardless of where all other particles are.

$$\Rightarrow f_\alpha = \langle F_\alpha \rangle = N_\alpha \int d\vec{X}_{\alpha_2} d\vec{X}_{\alpha_3} \dots d\vec{X}_{\alpha_{N_\alpha}} d\vec{X}_{\beta_1} \dots d\vec{X}_{\beta_{N_\beta}} \dots P_N$$

Similarly, two-particle dist. func:

$$f_{\alpha\beta} = N_\alpha N_\beta \int d\vec{X}_{\alpha_2} \dots d\vec{X}_{\alpha_{N_\alpha}} d\vec{X}_{\beta_2} \dots d\vec{X}_{\beta_{N_\beta}} \dots P_N$$

(explain)

$$\Rightarrow \frac{\delta F_\alpha}{\delta t} + \left\langle [F_\alpha, H_m] \right\rangle_{P_N} = 0 \quad \begin{aligned} & \rightarrow [f_\alpha, \langle H_m \rangle] + \left(\frac{\delta F_\alpha}{\delta t}, \delta H_m \right) \\ & F_\alpha - f_\alpha \quad H_m - \langle H_m \rangle \\ & \text{nonlinearity} \end{aligned}$$

$$\Rightarrow \frac{\partial f_\alpha}{\partial t} + [f_\alpha, \langle H_m \rangle] = - \underbrace{\langle [8F_\alpha, 8H_m] \rangle}_{\text{affected by correlations}}$$

1-part. dist. advected through phase space by mean field (still must justify what "mean" means...)

bet. $8F_\alpha$ and E&M fluctuations, both of which are associated w/ discrete nature of charges.

Take electrostatic interactions... can show

$$\begin{aligned} \frac{\partial f_\alpha}{\partial t} + \vec{v} \cdot \nabla f_\alpha + \frac{q_\alpha}{m_\alpha} \left[\vec{E} + \frac{v}{c} \times \vec{B} \right] \cdot \frac{\partial \vec{r}}{\partial t} \\ = - \sum_{\beta} q_\beta \int d\vec{x}' \underbrace{\frac{2}{\pi} \frac{q_\beta}{|\vec{r} - \vec{r}'|}}_{-\delta\varphi} \cdot \frac{\partial}{\partial \vec{r}} \left[f_\alpha(t, \vec{x}) f_\beta(t, \vec{x}') - f_{\alpha\beta}(t, \vec{x}, \vec{x}') \right] \end{aligned}$$

If all particles are completely uncorrelated, then $f_\alpha f_\beta = f_{\alpha\beta}$

\Rightarrow Vlasov (write it)

Write $f_{\alpha\beta} = f_\alpha f_\beta + g_{\alpha\beta}$ \curvearrowright 2-part. correlation function (explain)

$$\Rightarrow \boxed{\dot{f}_\alpha = \sum_f \frac{q_\alpha}{m_\alpha} \int d\vec{x}' \frac{2}{\pi} \frac{q_\beta}{|\vec{r} - \vec{r}'|} \cdot \frac{\partial}{\partial \vec{r}} g_{\alpha\beta}}$$

1st step of BBGKY hierarchy (Bogoliubov, Born, Green, Kirkwood, Yvon: 1935-1949)

Evolution of 1-part. DF depends on correlations b/w
2 parts. Relationships b/w discrete particles effect
future of each of them.

Can write eqn. for $g_{\alpha\beta}$:

$$\left(\frac{2}{\pi} + v \cdot \dot{v} + a \cdot \frac{2}{\pi v} + v' \cdot \dot{v}' + a' \cdot \frac{2}{\pi v'} \right) g_{\alpha\beta}$$

(cons. move through phase space, pushed by smooth fields)

$$= \sum_{\alpha} q_{\alpha} \int dx'' \frac{2}{\pi r} \frac{q_{\alpha}}{|r-r''|} \cdot \frac{\partial f_{\alpha}}{\partial v} g_{\alpha\beta}$$

(α forms relationships w/ the bath → will be Debye shielding
→ affects α relationship)

$$+ \sum_{\beta} q_{\beta} \int dx'' \frac{2}{\pi r'} \frac{q_{\beta}}{|r'-r''|} \cdot \frac{\partial f_{\beta}}{\partial v'} g_{\alpha\beta}$$

(β forms relationships w/ the bath → affects β relationship)

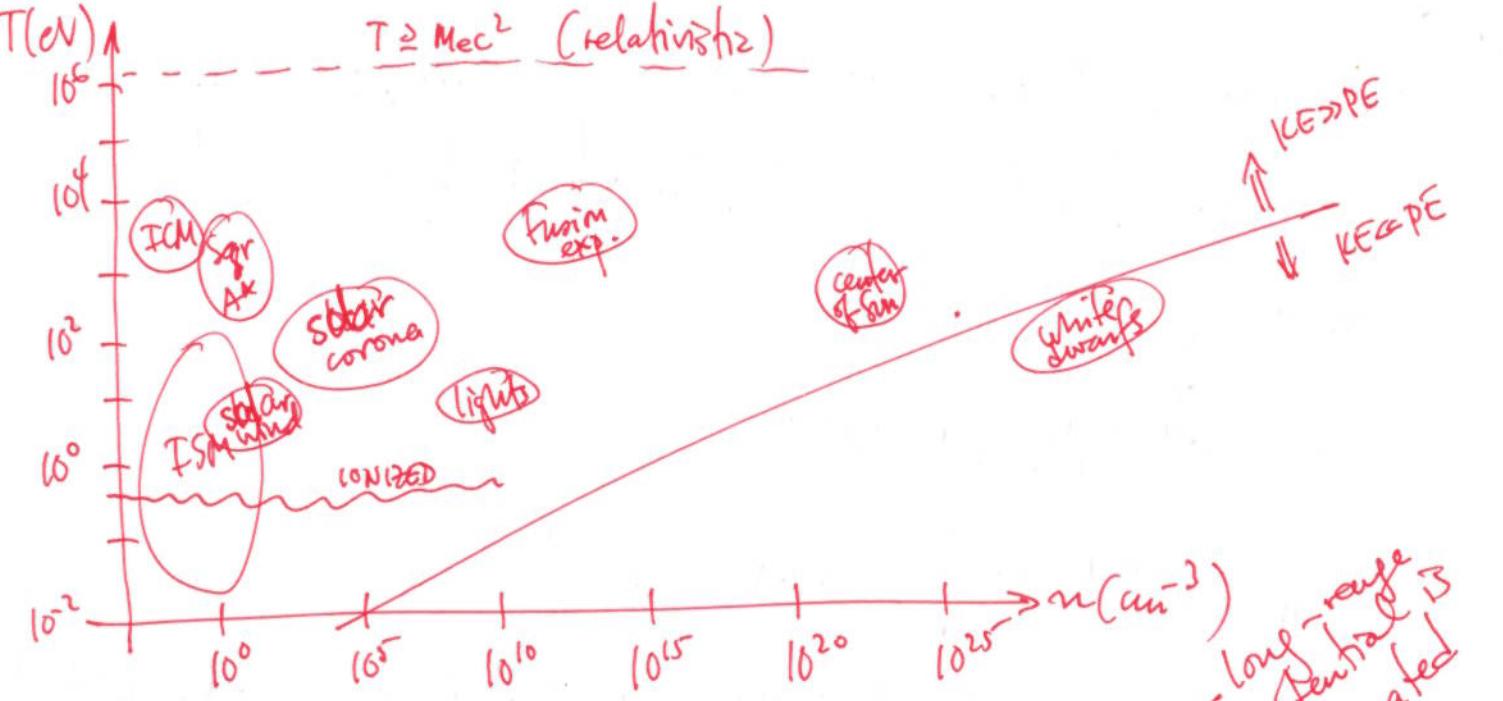
$$+ \frac{2}{\pi} \frac{q_{\alpha} q_{\beta}}{|r-r'|} \cdot \left(\frac{1}{m_{\alpha}} \frac{2}{\pi r} - \frac{1}{m_{\beta}} \frac{2}{\pi r'} \right) \left(f_{\alpha} f_{\beta} + \underbrace{q_{\alpha} q_{\beta}}_{\substack{\text{existing correlations} \\ \text{continue to source themselves.}}} \right)$$

f_{α}, f_{β} initially uncorrelated
parts form a relationship

$$+ \sum_{\gamma} q_{\gamma} \int dx'' \left(\frac{1}{m_{\alpha}} \frac{2}{\pi r} \frac{q_{\alpha}}{|r-r''|} \cdot \frac{2}{\pi r} + \frac{1}{m_{\beta}} \frac{2}{\pi r'} \frac{q_{\beta}}{|r'-r''|} \cdot \frac{2}{\pi r'} \right) g_{\alpha\beta}$$

(ménage à trois)

Keeps going... fortunately, straightforward to close
for a plasma...



When $KE \gg PE$, BBkry in equilibrium gives

$$f_{\alpha p} = f_{\alpha}(\vec{v}) f_p(\vec{v}) \left[1 - \underbrace{\frac{q_{\alpha} q_p}{T}}_{\text{Maxwellians}} \underbrace{\frac{e^{-r/\lambda_D}}{r}}_{\text{where } \lambda_D^2 = \left(\sum_{\alpha} \frac{4\pi q_{\alpha}^2 n_{\alpha}}{T} \right)^{-1}} \right] \sim \frac{1}{\lambda} e^{-x} \quad w/ \quad x = n \lambda_D^3 \ggg 1.$$

\rightarrow "quasi-neutral" on scales $\gg \lambda_D \approx 7 m \left(\text{TeV/cm}^{-3} \right)^{1/2}$. $\lambda \sim \frac{\text{dust}}{\text{Debye}}$

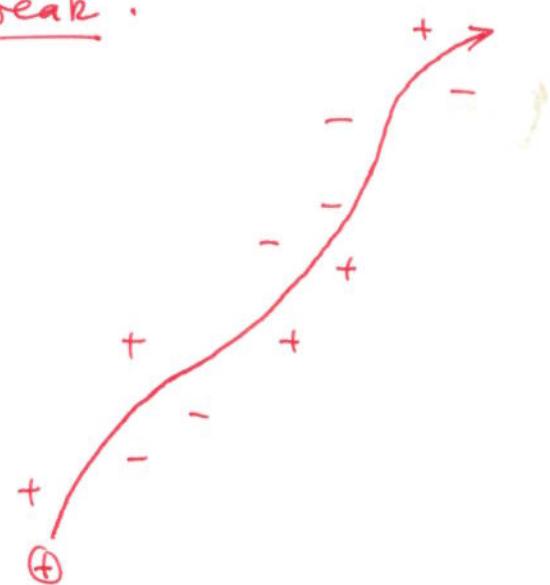
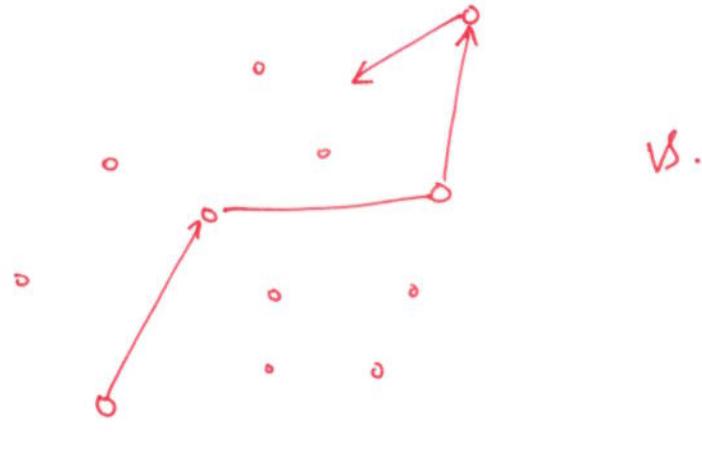
* tells us that correlations are weak for $\ell \gg \lambda_D$, and particles are almost distributed uniformly.

leads to "Bogoliubov's timescale hierarchy":

- On lengthscales $\sim \lambda_D$ and timescales $\sim 14 h^{-1} \lambda_D$, pair correlations are established and Coulomb potentials are shielded. (This is a reversible process.)

- On lengthscales $\sim \lambda_{\text{Df}} \sim \sqrt{k_B T} / v \sim \lambda_D \left(\frac{\Lambda}{\ln \Lambda} \right)$ and timescales $\sim v^{-1} \sim \omega_p^{-1} \left(\frac{\Lambda}{\ln \Lambda} \right)$, local Maxwellian VDFs are attained via collisional relaxation.
- Macroscopic dynamics on lengthscales $\sim L$ and timescales $\sim U^{-1} L \sim \left(\frac{\sqrt{k_B T}}{U} \right) v^{-1} \left(\frac{L}{\lambda_{\text{Df}}} \right) \gg v^{-1}$ or $\gtrsim v^{-1}$
 "collisional" "weakly
 (fluid) "collisional"
 or "collisionless"

This is completely different from the situation of a gas of neutral particles. There, the interaction range is $\propto r_0 \ll \text{mean}$ spacing $\sim n^{-1/3} \Rightarrow nr_0^3 \ll 1 \Rightarrow$ collisions are rare. By contrast, in a plasma, $r_0 \approx \lambda_D \gg n^{-1/3} \Rightarrow nr_0^3 \gg 1 \Rightarrow$ collisions (actually, correlations) happen all the time... NOT rare. Even though a charged particle is interacting w/ all particles in its Debye sphere and thus undergoes $\sim \Lambda$ simultaneous interactions, most interactions are weak.



- ↗ straight-line trajectories
- ① Assume f_α is homogeneous on scale over which q_α varies.
 - ② Assume particles enter into relationships initially unacquainted (Markov assumption) (good b/c $(\gg 1)$)
 - ③ Drop 3-body correlations (weak—not rare)

$$\Rightarrow \left(\frac{\partial f_\alpha}{\partial t} \right)_{\text{Vlasov}} = \bar{v}_f \int \frac{d\vec{k}}{(2\pi)^3} \frac{\vec{k}}{m_\alpha} \cdot \frac{\partial}{\partial \vec{v}} \int d\vec{v}' \left(\frac{4\pi q_\alpha q_f}{k^2 D(\vec{k}, \vec{v}, \vec{v}')} \right)^2$$

$$x \pi \delta(\vec{k}\vec{v} - \vec{k}\vec{v}') \left(\frac{\vec{k}}{m_\alpha} \cdot \frac{\partial}{\partial \vec{v}} - \frac{\vec{k}}{m_f} \cdot \frac{\partial}{\partial \vec{v}'} \right) f_\alpha(\vec{v}) f_f(\vec{v}')$$

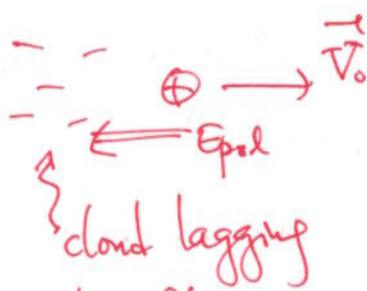
$$= \left(\frac{\partial f_\alpha}{\partial t} \right)_{c, B-L}$$

w - $\vec{k} \cdot \vec{v} = 0$

2 initially uncorrelated particles...
... push equal and oppositely on one another in direction of their Coulomb electric fields...
... interaction requires resonance (Landau-resonant excitation of waves, Landau-resonant absorption of waves)

Dressed potential: Coulomb $\propto \frac{1}{k^2}$
shielding by equilibrated Debye cloud in frame of particle

Explain polarization drag via pair correlation est'd on a w_p time



Decel. Deceleration sources waves... these contribute to "thermal bath" off of which other particles can diffuse)

New Debye cloud after every deflection of charge

Because $\gg \omega_p$, charges always carry well-equilibrated Debye clouds w/ them.

Rostoker test-particle superposition principle: plasma composed of uncorrelated Debye clouds. \Rightarrow Poisson-Lenard

In equilibrium, minimum level of Electron noise given by

$$W = \int \frac{dk}{(2\pi)^3} \frac{1}{2} \frac{1}{1 + (k/k_0)^2}$$

$T/2$ per degree of freedom; fluctuation-dissipation relation

$$(P_{w,k} = -2\omega W_{w,k} \tan[D(\omega_k)])$$

power lost by polarization drag = power in emitted and Landau-damped waves, which are absorbed by other particles resonant

BL: $\frac{\partial f_\alpha}{\partial t} = - \frac{2}{n} \cdot \left[\overbrace{\tilde{A} f_\alpha}^{\left\{ \begin{array}{l} q_x \rightarrow E_{\text{pol}} \\ n_{\text{tot}} \end{array} \right\}} + \frac{1}{2} \frac{\partial}{\partial v} \frac{\partial}{\partial v} : \left[\overbrace{\tilde{B} f_\alpha}^{\left(\begin{array}{l} q_x \\ \text{from imperfect} \\ \text{Debye shielding} \end{array} \right)} \right] \right]$

$\left(\frac{q_x}{n_{\text{tot}}} \right)^2 \langle \tilde{E} \tilde{E}^\dagger \rangle_{\omega = k \cdot v}$
 (from resonant diffusion off quasi-thermal bath)

PP form
 This confirms that scatterings are small angle.
 OR
 $- \frac{2}{n} \cdot \left[\frac{q_x}{n_{\text{tot}}} \tilde{E}_{\text{pol}} \right] + \frac{1}{2} \frac{\partial}{\partial v} \left[\overbrace{\tilde{B} \cdot \frac{\partial f_\alpha}{\partial v}}^{\text{calculated}} \right]$

Too complicated to implement in plasma^{calculator} — no one ever does this. But it does spell out the path forward for understanding other stochastic processes, e.g., diffusion/drag in a spectrum of waves ("quasi-longer-scale"), (linear operator), and it is useful for computing spatial diffusion coefficients related to collisional relaxation of temperature and flow gradients (Chapman-Enskog-Braginskii).